

Poynting Vector, Oscillating charges, Polarization

Electric field density energy

$$U_E = \frac{1}{2} \epsilon_0 E^2 \quad [\text{J/m}^3]$$

Magnetic field density energy

$$U_B = \frac{1}{2\mu_0} B^2 \quad [\text{J/m}^3]$$

$$= \frac{1}{2\mu_0} \frac{E^2}{c^2} = \frac{1}{2} \epsilon_0 E^2$$

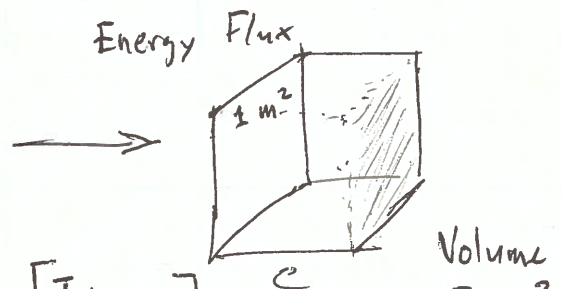
Total Energy density

$$U_{\text{total}} = U_E + U_B = \epsilon_0 E^2 = \epsilon_0 E B c$$

Energy flux density...

$$U_{\text{total}} c = \epsilon_0 E B c^2 = \frac{E B}{\mu_0} \quad [\text{J/m}^2 \text{sec}]$$

↑ Poynting vector

POYNTING VECTOR

$$\vec{S} = \vec{E} \times \vec{B} / \mu_0 \quad \left[\frac{\text{W}}{\text{m}^2} \right]$$

"the energy transfer per unit area per unit time"

"Energy Flux $\vec{S} = \vec{E} \times \vec{H}$ "

Time Averaged Value of Poynting Vector

$$\langle S \rangle = \frac{1}{2} E_0 B_0 / \mu_0 = \frac{1}{2} \frac{E_0^2}{\mu_0 c}$$

"the instantaneous power flow due to instantaneous E/M fields"

Example

Plane E/M wave

$$E_0 = 100 \text{ V/m}$$

$$\langle S \rangle = \frac{1}{2} \frac{100^2}{\mu_0 c} = 13 \text{ W/m}^2$$

(Poynting vector independent of frequency f, ω)

$$\text{Now, } E_0 = 1,000 \text{ V/m}$$

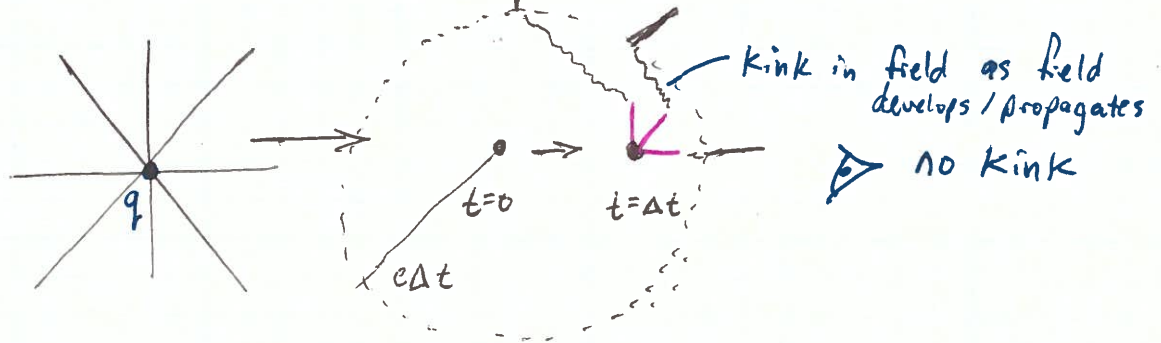
$$\text{so... } \langle S \rangle = 1.3 \text{ kW/m}^2$$

note S scales as E^2

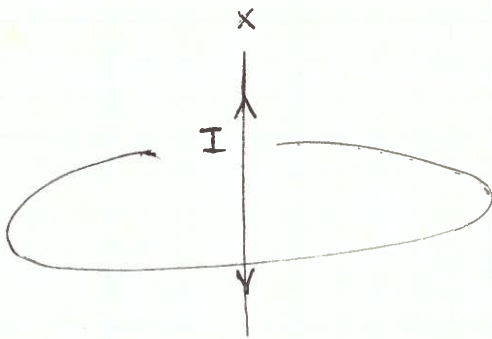
Note, plane waves exist at all time in all of space
(planes are infinite in math...)

How is an EM wave produced?

Begin w/ stationary charge, accelerate for time Δt

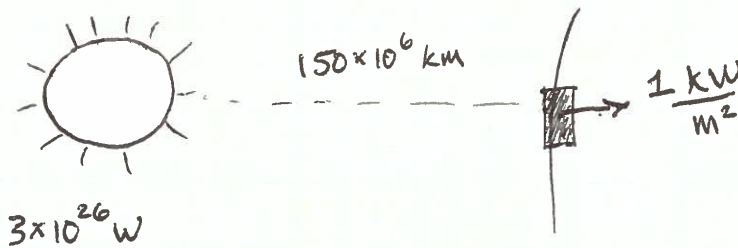


Charges on a wire, antenna



Plane wave solution is not very realistic

Sun's energy at Earth, Poynting vector and E-Field



$$\langle S \rangle = \frac{1}{2} \frac{E_0^2}{\mu_0 c}$$

$$\downarrow$$

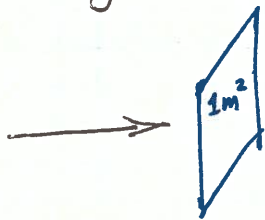
$$1,000 = \frac{1}{2} E_0^2 / \mu_0 c$$

$$\underline{E_0 \approx 870 \text{ V/m}}$$

Photons are individual packets (\sim bullets) that carry energy and so it has momentum.

$$p = \frac{\text{energy of 1 photon}}{c}$$

Poynting Vector and Momentum



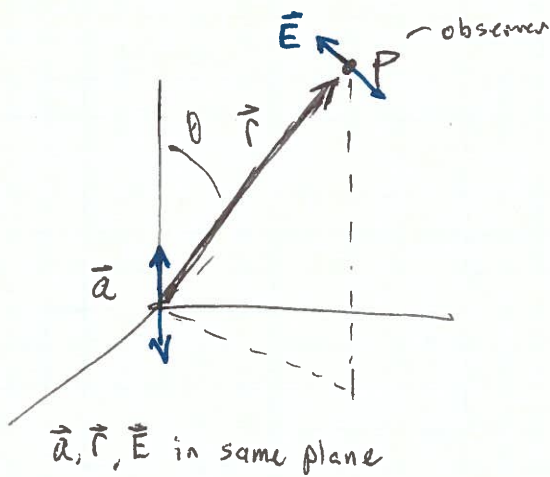
$$\frac{S}{c} = \frac{\text{Energy}}{c \cdot \text{m}^2 \cdot \text{sec}} = \text{Pressure}$$

$$\text{So... } \frac{\langle S \rangle}{c} \propto = \text{Radiation Pressure}$$

Comets have 2 tails...

one is white due to sun's radiation pressure

other is blue due to solar wind



Oscillating charge is accelerated

We know, $\vec{E} \perp \vec{r}$, $E \propto q \frac{a \sin \theta}{r}$

$$\text{So, } S \propto q^2 \frac{a^2}{r^2} \sin^2 \theta$$

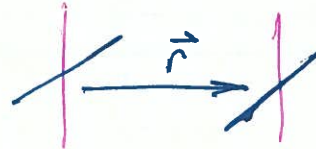
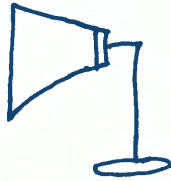
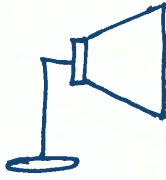
↑
Poynting Vector

(Conservation of energy)

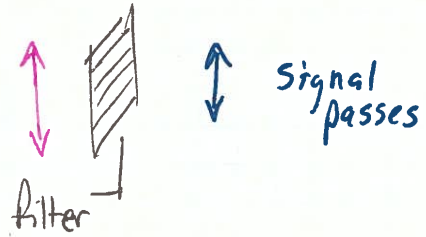
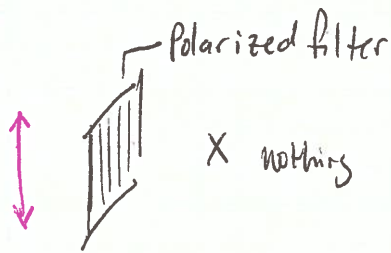
DEMO Linear Polarization

transmitter transmits at 10 GHz ($\lambda = 3\text{cm}$)

audio modulated $\sim 1\text{kHz}$



antenna + receiver orientation must match to work!



DEMO

75 MHz ($\lambda = 4\text{m}$)

